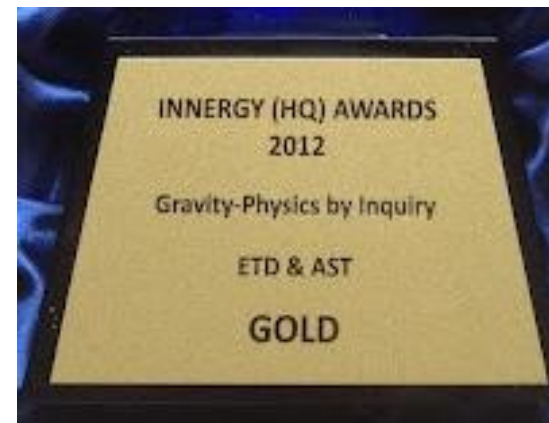


# Easy Java Simulation (EJS), an innovative tool for teacher as designers of gravity-physics computer models



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- F. Esquembre, F-K Hwang, W. Christian, D. Brown, M. Belloni, A. Cox, A. Duffy, T. Mzoughi, M. Gallis, T. Timberlake, J. M. Aguirregabiria, W. Junkin, H. Gould, J. Tobochnik, Jose Sanchez, S. Tuleja and many more....

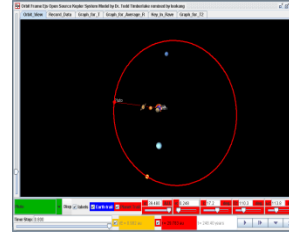
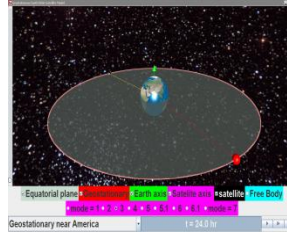
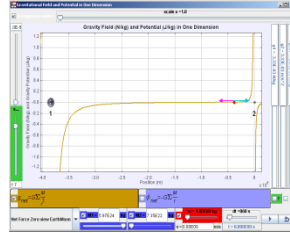
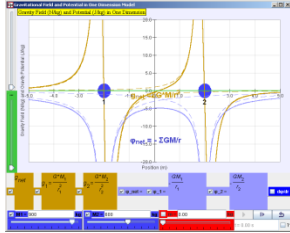
## • Digital Libraries

- <http://www.compadre.org/OSP/>
- <http://www.phy.ntnu.edu.tw/ntnujava/index.php>
- EJS itself has examples as well



# 1. innovativeness of EJS

## • Breakthrough in teacher education [R]



### – products (computer models)

- **Customized** to SG syllabus (**teacher-student feedback**)

### – processes ( professional community)

- **Teacher education in OSP community**
- To **benefit** all **humankind**

## 2. Rationale

- Physics is best learnt **by inquiry** through **hands-on** exploration.
- Difficult to conduct experiment **[R]** on topics like Gravitation in school laboratories.
- Many students **fail to visualise** the gravitational effects of huge masses and the motion of satellites.
- **Little physical meaning** to the number-crunching.



# 3. Design

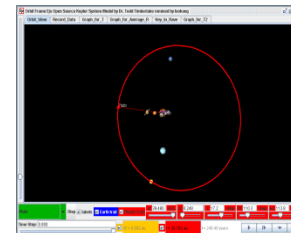
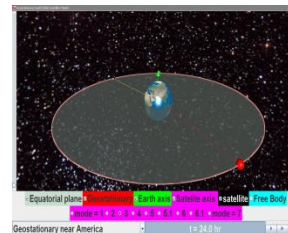
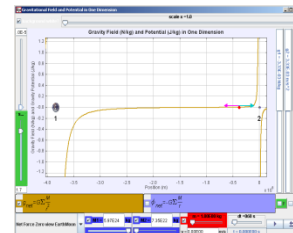
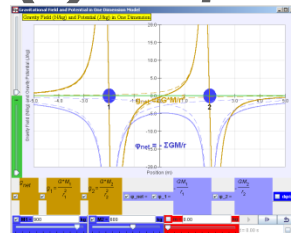
- **Four** simulations were selected to address concepts on:

*(1) Gravitational field strength and potential;*

*(2) Escape velocity;*

*(3) Geostationary orbit;*

*(4) Kepler's Third law.*



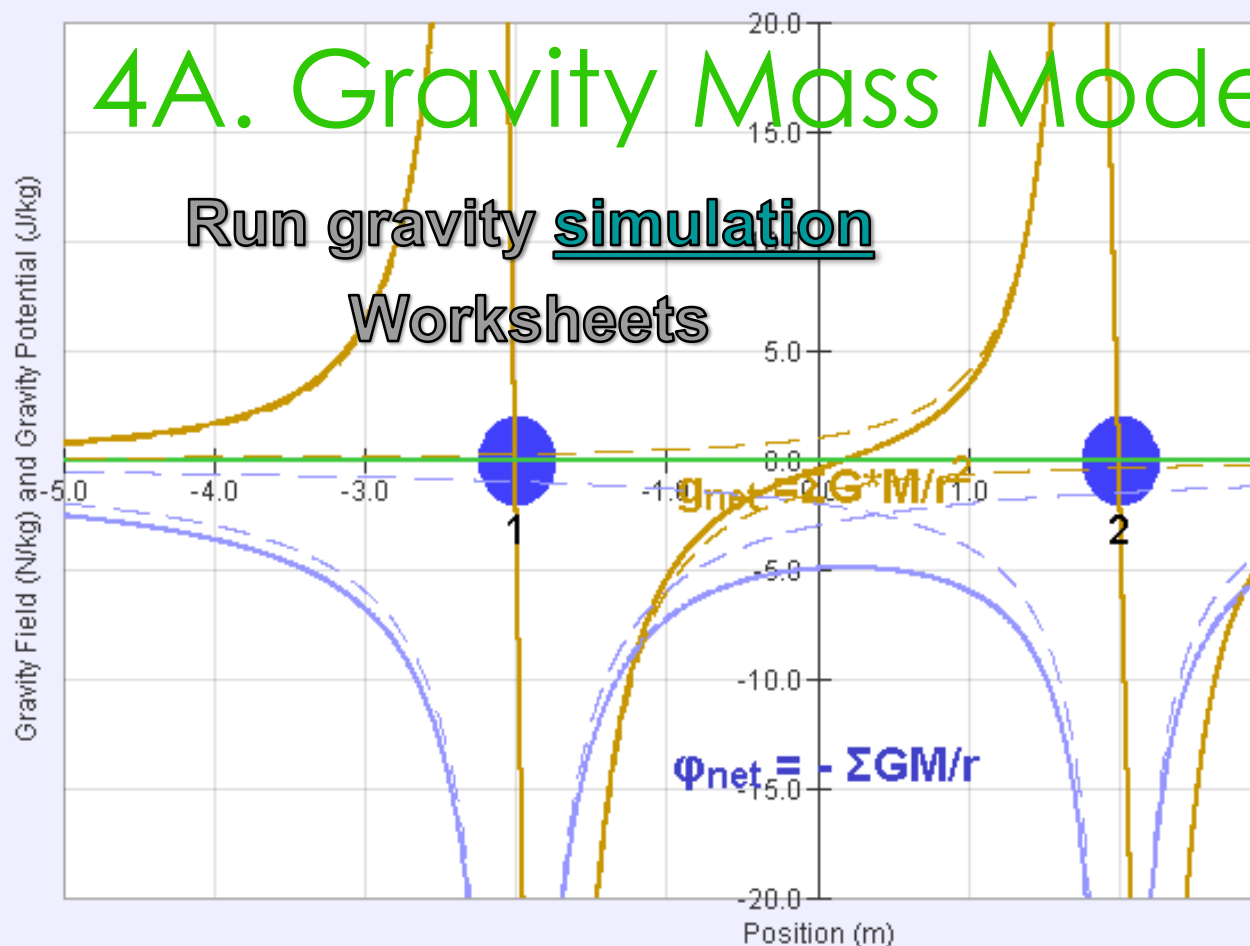
- These simulations were infused strategically into **tutorial** questions or separately as **laboratory worksheets** to enhance students' learning.

### 3. Design (modified 5E inquiry)

- Students were introduced to relevant videos and get them **e**ngaged in the simulations.
- Students **e**xplored the simulations at their own pace, guided by the worksheet questions. ***[Self-directed learning]***
- Students **e**xplained their understanding in small groups. ***[Collaborative learning]***
- Teacher asked questions to get students to **e**laborate their understanding.
- Selected groups did class presentation and **e**valuate what they have learnt.

# 4A. Gravity Mass Model

Run gravity simulation  
Worksheets



- Near 'impossible' to experience gravity due to small mass on Earth

- Invisible concepts made alive and experience interconnected

- Data analysis

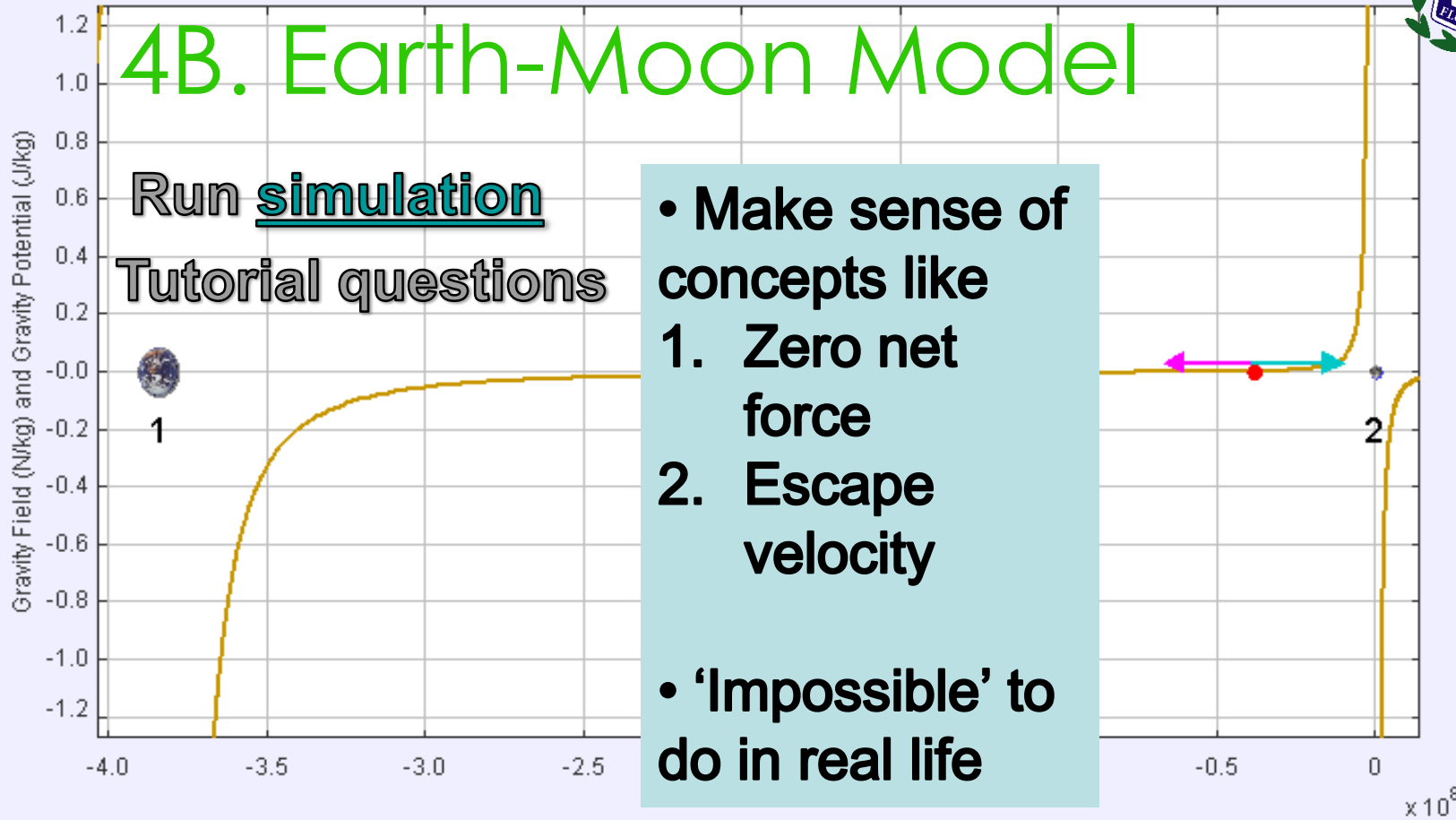
☒  $g_{net}$ 
☒  $g_1 = \frac{G \cdot M_1}{r_1^2}$ 
☒  $g_2 = \frac{G \cdot M_2}{r_2^2}$ 
☒  $\phi_{net} =$ 
☒  $\phi_1 =$ 
☒  $\phi_2 =$

☒  $M_1 = 900$  kg
 ☒  $M_2 = 600$  kg
 ☐  $m = 0.00$  kg

# 4B. Earth-Moon Model

Run simulation  
Tutorial questions

- Make sense of concepts like
- 1. Zero net force
- 2. Escape velocity
- 'Impossible' to do in real life



☒  $g_{\text{net}} = G \sum \frac{M}{r^2}$  ☐  $\phi_{\text{net}} = -G \sum \frac{M}{r}$  ☐ F ☐ U

Net Force Zero view EarthMoon ☒ M1 = 5.97E24 kg ☒ M2 = 7.35E22 kg ☒ m = 1.00E00 kg dt = 060 s

v = 0.00E00 m/s t = 0.000E00 s

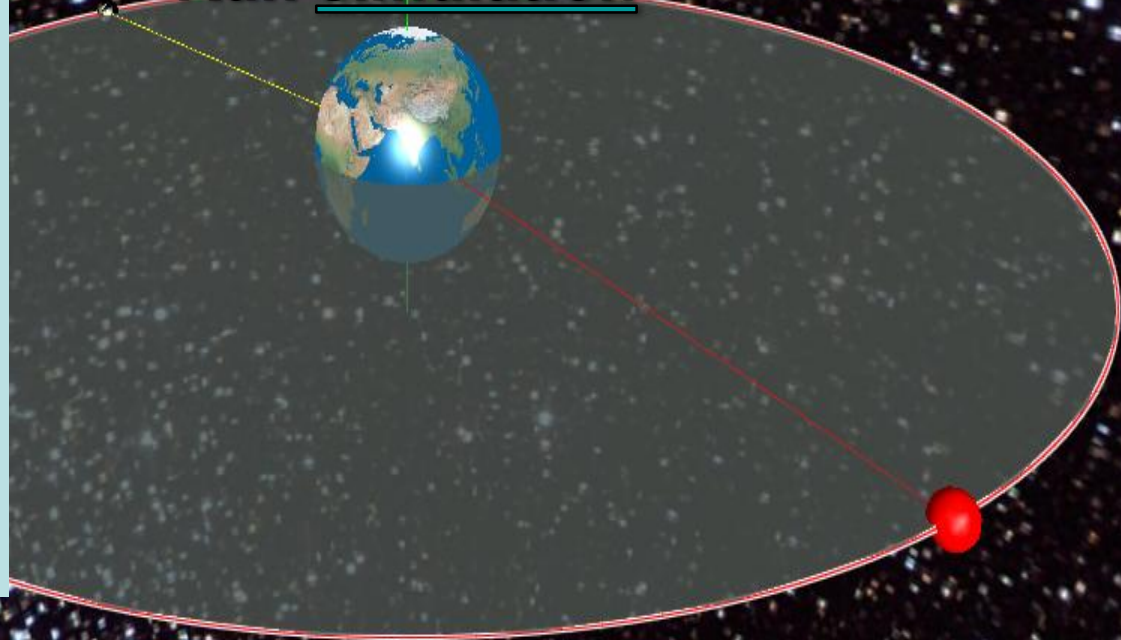




# 4C. Geostationary Satellite Model

- Teacher is not the only source of knowledge
- Learning is experiential & involves meaning making through visualization

Run simulation



☒ Equatorial plane ☒ Geostationary ☒ Earth axis ☒ Satellite axis ☒ satellite ☐ Free Body

• mode = 1 • 2 • 3 • 4 • 5 • 5.1 • 6 • 6.1 • mode = 7

Geostationary near America

t = 24.0 hr



[Derived work based on Francisco \(Paco\) Esquembre's earlier model](#)

# 4C. Geostationary Satellite Model

- Create active learning previously not possible

The students' answers are now actually based on the acts of scientific inquiry being scientist themselves, rather than the memorization of facts (traditional teaching method).

Making logical conclusions based on student-lead evidence based inquiry on the models.

Repeat step 8 for mode 2 & 3. Mode 2 shows a geostationary satellite moving above Africa, while mode 3 shows a geostationary satellite moving above America.

Conclusion about the relative positions of the satellite and Africa/America: Both are in

the same rotation. Satellite is always vertically above Africa and America if it rotates around the Earth.

Determine the period of rotation of the Earth about its own axis: 24 hours

Determine the period of rotation of the geostationary satellite: 24 hours

What is the satellite's radius of rotation, from the centre of the Earth? 35700 km

10) Hence why do you think the satellite is called a "geostationary satellite", when the satellite is obviously not stationary? The satellite always stays above a fixed point

as it rotates around the Earth. Hence it looks stationary from the fixed point's perspective.

Orbit\_View

Record\_Data

Graph\_for\_T

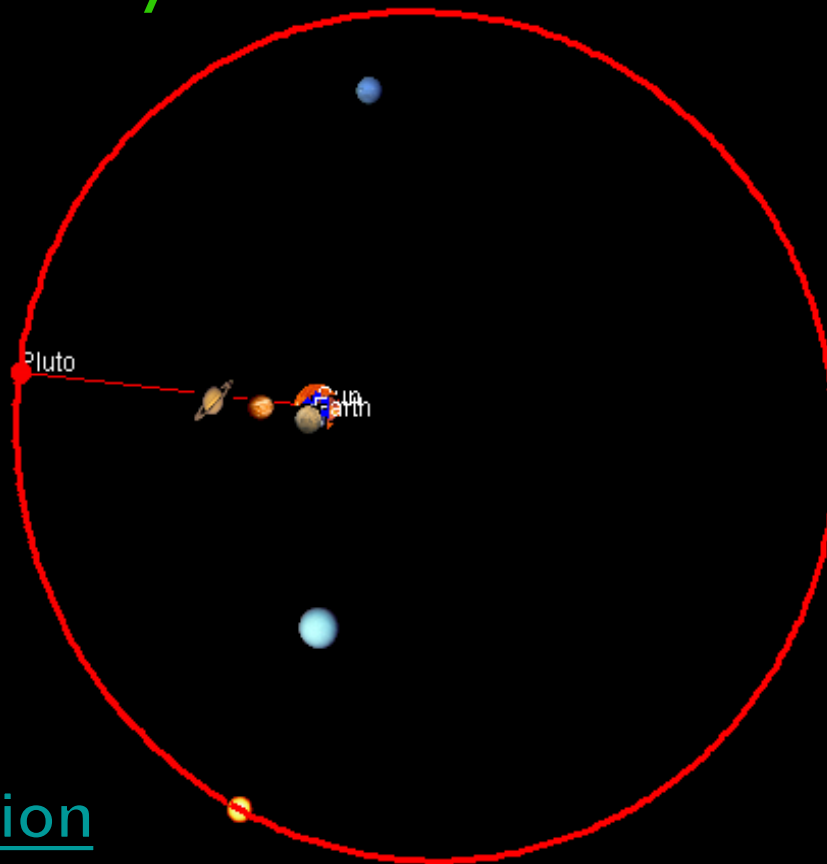
Graph\_for\_Average\_R

Key\_In\_Rave

Graph\_for\_T2

# 4D. Kepler's System Model

- Based on real data
- Impossible to do in real life



Run [simulation](#)

Pluto

Disp

☒ labels☒ Earth trail☒ Planet Trail

a: 39.480

A.U.

e: 0.248

i: 17.2

deg.

Q: 110.3

deg.

ω: 113.8

deg.

Time Step: 0.800



rE = 0.992 au



r = 29.783 au

t = 248.40 years





Orbit_View	Record_Data	Graph_for_T	Graph_for_Average_R	Key_In_Rave	Graph_for_T2
#	T / years	R / A.U.	Raverage / A.U.		
0	0.245	0.387	0.000		
1	0.615	0.723	0.000		
2	1.000	1.000	0.000		
3	1.880	1.524	0.000		
4	11.888	5.203	0.000		
5	29.500	9.537	0.000		
6	84.200	19.190	0.000		
7	164.000	30.060	0.000		
8	248.400	39.480	0.000		

- data collect inside EJS instead of pen paper

$T = Ar^n$   
 $\ln T = \ln(Ar^n)$

$rE = 0.999 \text{ au}$

Planet	$\tau$ / year	$r / 10^6 \text{ km}$	$\ln T$	$\ln r$
Mercury	0.24 <sup>-1.43</sup>	67.3	-1.43	18.0
Venus	0.62 <sup>-0.47</sup>	109	-0.47	18.5
Mars	1.88 <sup>0.63</sup>	210	0.63	19.2
Jupiter	11.87 <sup>2.47</sup>	799	2.47	20.5
Saturn	29.45 <sup>3.38</sup>	1469	3.38	21.1

Pluto

Disp

☒ labels☒ Earth trail☒ Planet Trail

a: 39.480 A.U.

e: 0.248

i: 17.2 deg.

Q: 110.3 deg.

w: 113.8 deg.

Time Step: 0.800

☒  $rE = 0.992 \text{ au}$ ☒  $r = 29.783 \text{ au}$ 

t = 248.40 years



Orbit\_View

Record\_Data

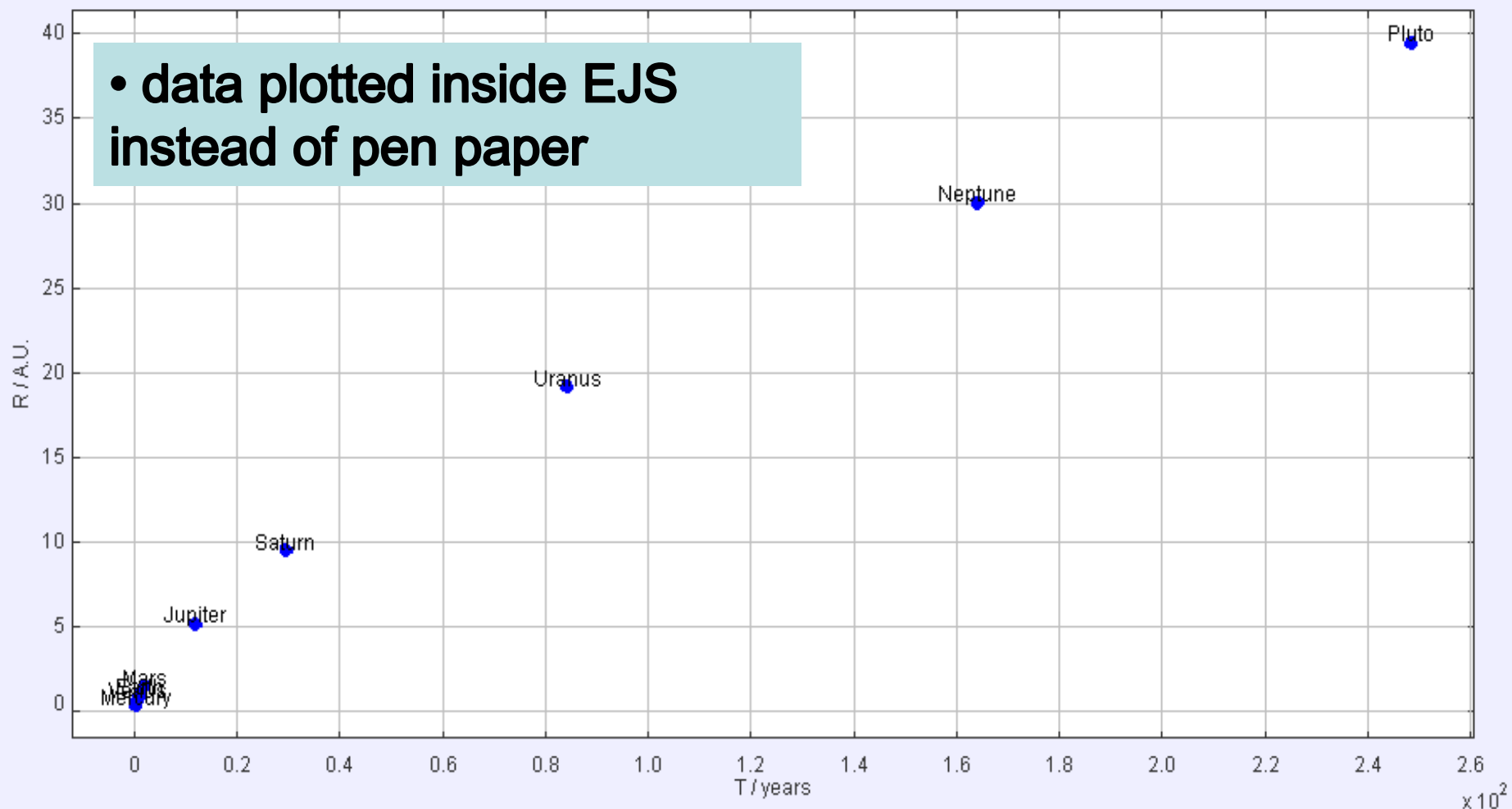
Graph\_for\_T

Graph\_for\_Average\_R

Key\_In\_Rave

Graph\_for\_T2

R versus T



Pluto

Disp ☒ labels☒ Earth trail☒ Planet Trail

a: 39.480 A.U.

e: 0.248

i: 17.2 deg.

Q: 110.3 deg.

 $\omega$ : 113.8 deg.

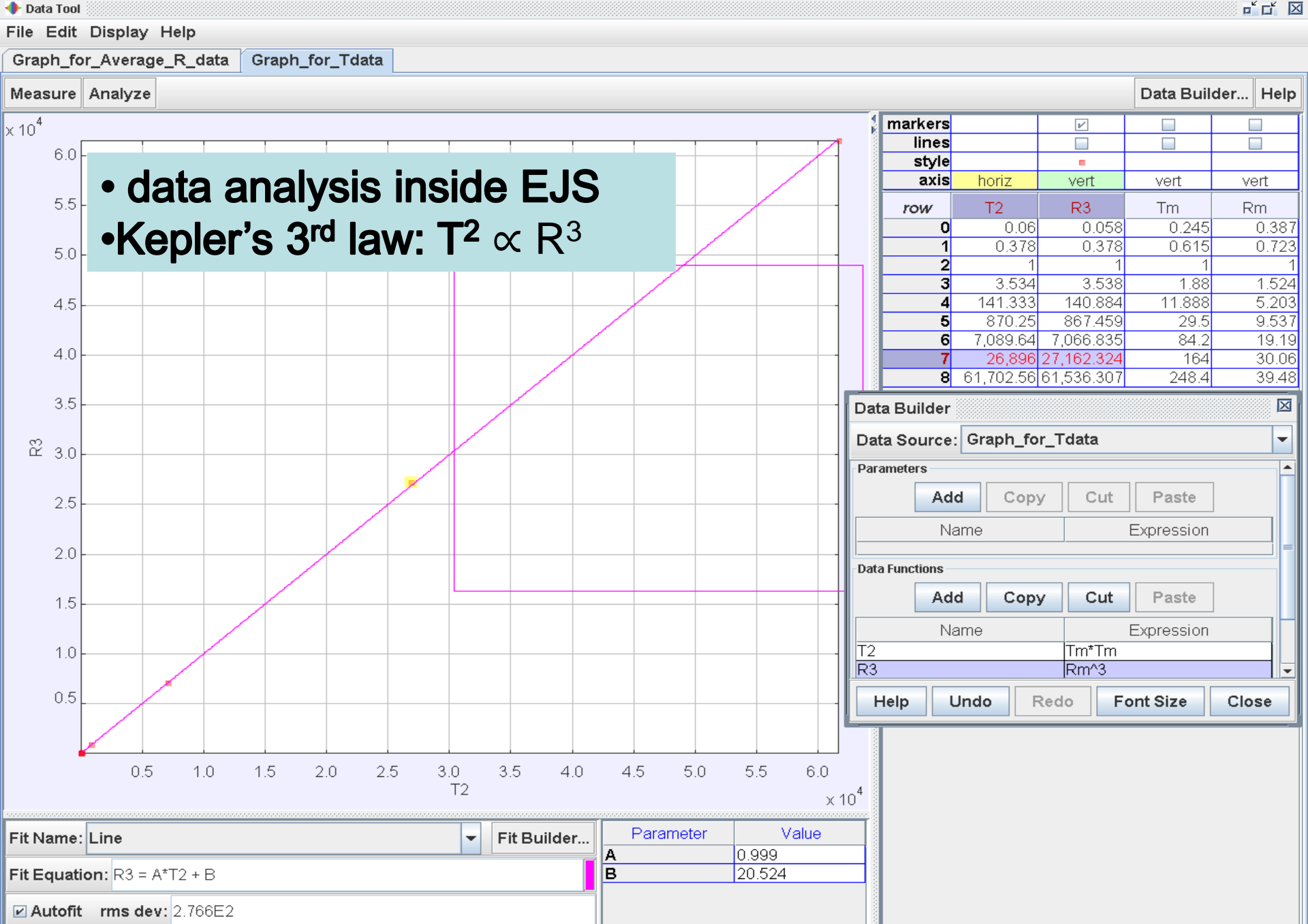
Time Step: 0.800

☒ rE = 0.992 au☒ r = 29.783 au

t = 248.40 years







Derived work based on Todd Timberlake's earlier model

- data analysis inside EJS
- Kepler's 3<sup>rd</sup> law:  $\log T$  vs  $\log R$  with evidence of relationship  $A=1.5$



Fit Name: Line Fit Builder...

Fit Equation:  $\ln T = A \ln R + B$

☒ Autofit rms dev: 2.125E-3

Parameter	Value
A	1.498
B	0.002

markers	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
lines	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
style	<input checked="" type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>	<input type="checkbox"/>
axis	horiz	vert	vert	vert	vert	vert
row	lnR	lnT	T2	Tm	R3	Rm
0	-0.412	-0.611	0.06	0.245	0.058	0.387
1	-0.141	-0.211	0.378	0.615	0.378	0.723
2	0	0	1	1	1	1
3	0.183	0.274	3.534	1.88	3.538	1.524
4	0.716	1.075	141.333	11.888	140.884	5.203
5	0.979	1.47	870.25	29.5	867.459	9.537
6	1.283	1.925	7,089.64	84.2	7,066....	19.19
7	1.478	2.215	26,896	164	27,162...	30.06
8	1.596	2.395	61,702...	248.4	61,536...	39.48

Data Builder

Data Source: Graph\_for\_Tdata

Data Functions

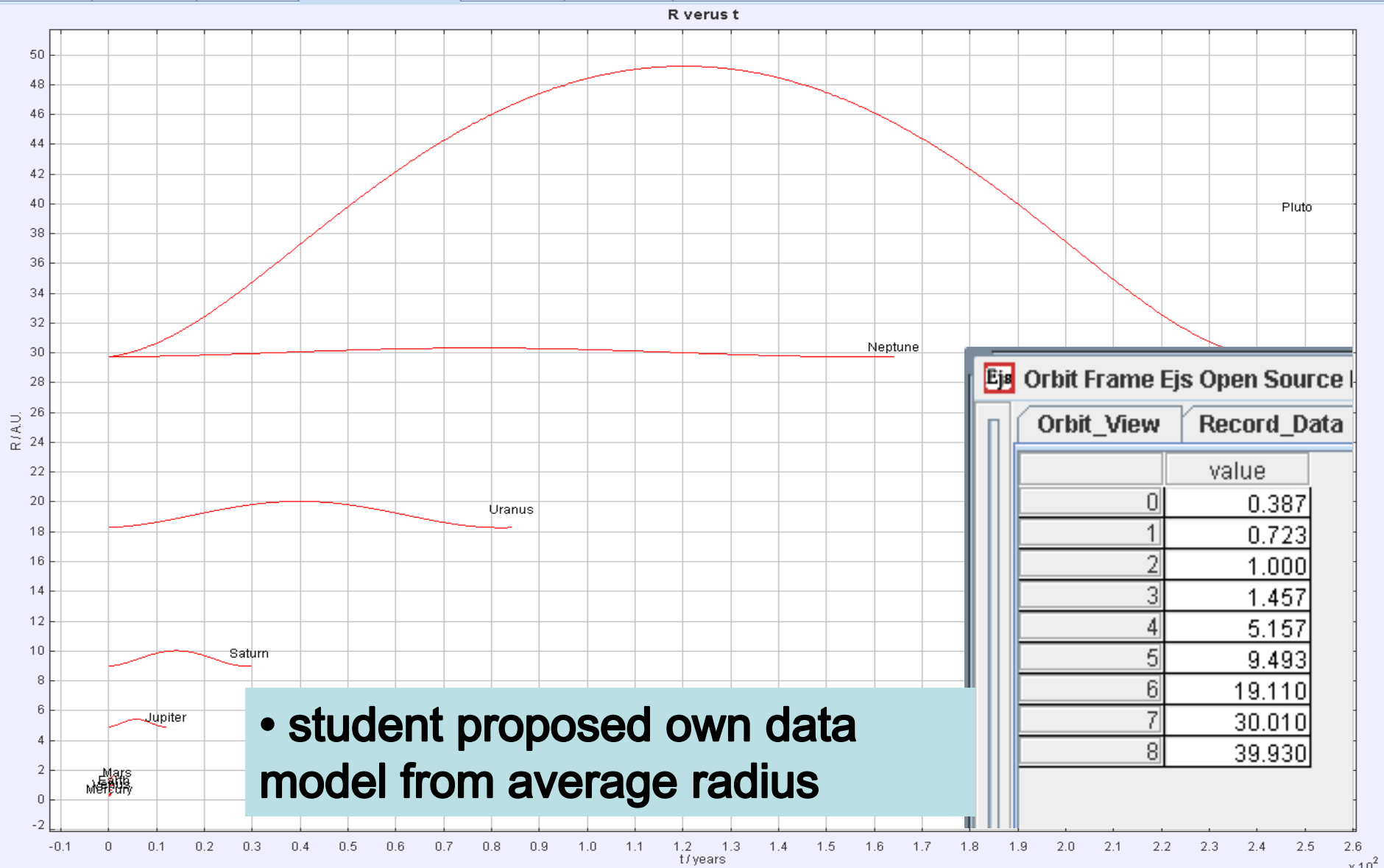
Add Copy Cut Paste

Name	Expression
T2	Tm*Tm
R3	Rm^3
lnT	log(Tm)
lnR	log(Rm)

Double-click cell to edit. See Help for valid expressions.

Help Undo Redo Font Size Close

Orbit\_View Record\_Data Graph\_for\_T Graph\_for\_Average\_R Key\_In\_Rave Graph\_for\_T2



Ejs Orbit Frame Ejs Open Source I

Orbit\_View

Record\_Data

	value
0	0.387
1	0.723
2	1.000
3	1.457
4	5.157
5	9.493
6	19.110
7	30.010
8	39.930

Pluto Display

☒ labels

☒ Earth trail

☒ Planet Trail

a: 39.480

A.U.

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deg.

Time Step: 0.800

☒ E = 0.992 au

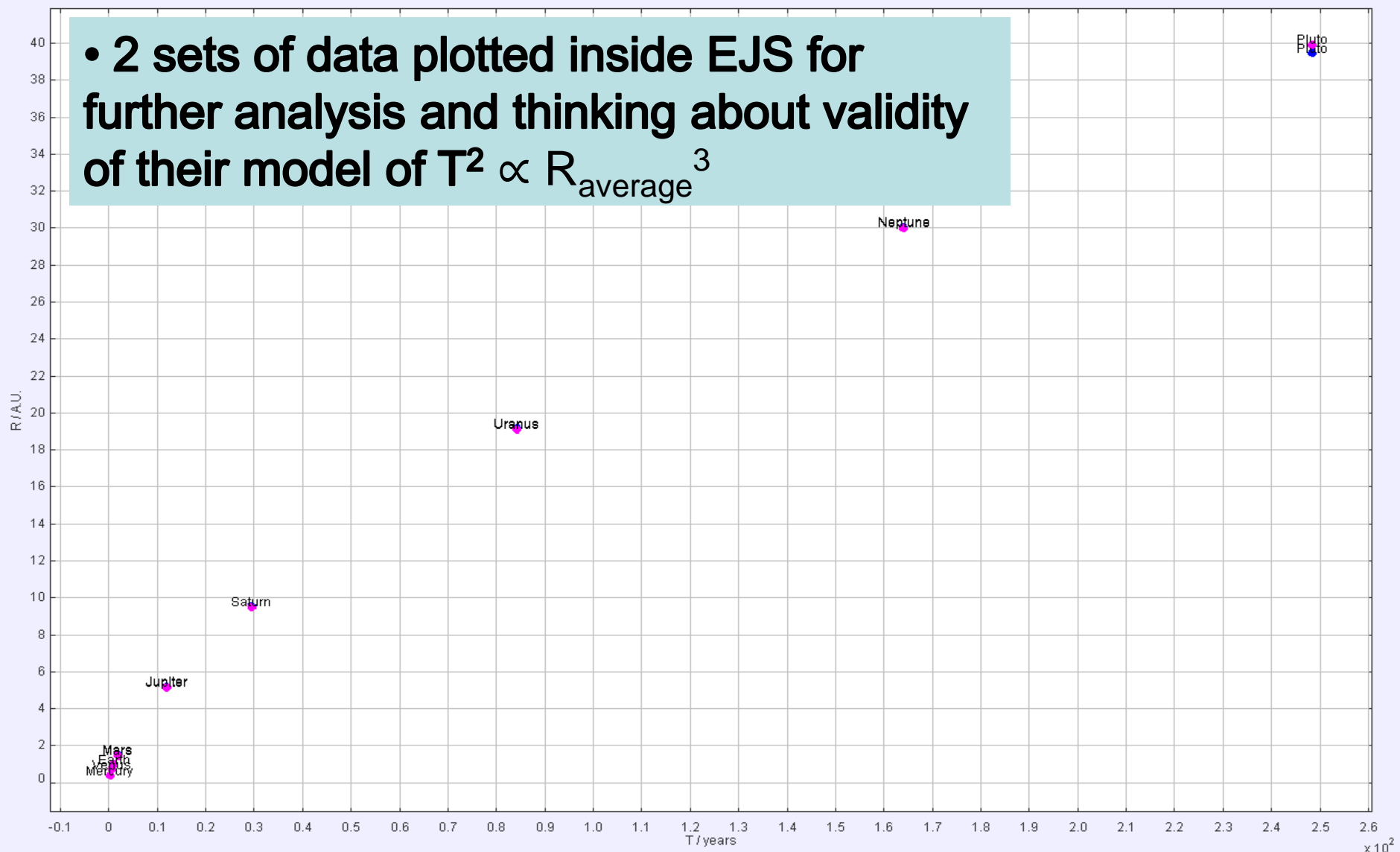
☒ r = 29.783 au

t = 248.40 years

Orbit\_View Record\_Data Graph\_for\_R\_vs\_T Graph\_for\_Average\_R\_vs\_T Key\_In\_Rave Graph\_for\_R\_combined\_vs\_T

R versus T

• 2 sets of data plotted inside EJS for further analysis and thinking about validity of their model of  $T^2 \propto R_{\text{average}}^3$



Pluto ☐ Display ☒ labels ☒ Earth trail ☒ Planet Trail

Time Step: 0.000

# 5A: Student feedback

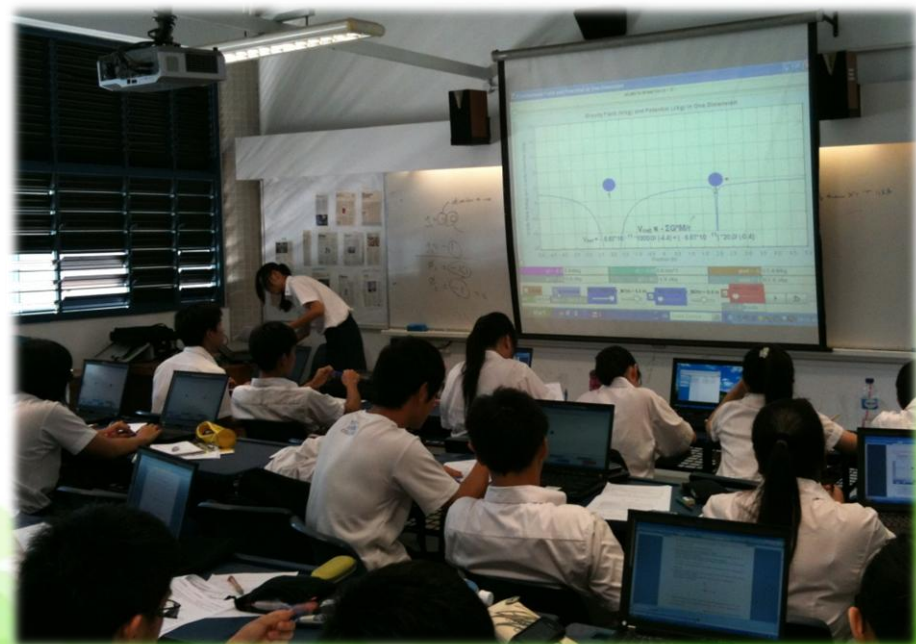
- The EJS lessons make it easier for us to depict the **motion of objects in 3D**.
- The theory is now much **easier to understand**.
- The simulations provided interesting information **beyond the syllabus**.
- “I think the simulations are good as they make me **view physics in a different perspective**.”





# 5B: Teacher reflection

- Help students **visualise** the effect of gravitation, motion of satellite, etc.
- Manipulation of parameters allows **independent learning** outside curriculum time.
- Curriculum time is required to **guide** students on the use of simulations at the initial stage.
- Worksheets questions need to be **more thought-provoking**.



# 6. we, the teachers are evidences that EJS has been an innovative tool

- **Learners as “scientists”**
  - Inquiry-based learning into the hands of ordinary students all over the world.

- **Teachers as designers** of computer models

- 5 journal **publication** at Physics Education
- 3 SG **local** conference papers
  - Academy Sym2011, iCTLT2012, ICCE2012

**IOP**  
Institute of Physics



- 4 **overseas** conference papers

- AAPTSM10 2010, AAPTWM12 2012, WCPE 2012, MPTL18

